

Relativistic Field Theory

-) When we make the transition from special relativity (SR) to GR, the Minkowski metric $\eta_{\mu\nu}$ will be promoted to a dynamical tensor field $g_{\mu\nu}(x)$.
-) GR is a classical field theory for the metric field $g_{\mu\nu}(x)$ whose dynamics is governed by the Einstein equations.
-) The Einstein equations can be derived from an action principle from the so-called Einstein-Hilbert action, which we will do later in the course.
-) Before we apply this to GR it is useful to remind ourselves how the action formalism works.
-) Action of a single point particle in classical mechanics:
 action $S[q] = \int dt L(q, \dot{q})$, $L = K - V$
 (Lagrangian) (kinetic energy) (potential energy)
-) The action is a functional, i.e., it maps a function $q(t)$ to a number $S[q]$.

path for $\delta S = 0$.

fixed

fixed

P_1

P_2

paths for $\delta S = 0$

$\delta S \neq 0$

-) The equations of motion of a system are derived from its action by demanding the principle of least action, also called variational principle:

A path taken by a physical system between two points is the one that extremizes the action:

$$\delta S = 0.$$

-) Let's now apply the variational principle to the action $S[q]$:

$$\begin{aligned}
 0 = \delta S &= \int_{t_1}^{t_2} dt \left(\frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \delta \dot{q} \right), \quad \underbrace{\delta q \Big|_{t_1} = \delta q \Big|_{t_2} = 0}_{\text{fixed boundary conditions.}} \\
 &= \int_{t_1}^{t_2} dt \left(\frac{\partial L}{\partial q} \delta q + \frac{\partial L}{\partial \dot{q}} \frac{d}{dt} \delta q \right) \\
 &= \int_{t_1}^{t_2} dt \left(\frac{\partial L}{\partial q} \delta q + \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \delta q \right) - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) \delta q \right) \\
 &= \int_{t_1}^{t_2} dt \left(\frac{\partial L}{\partial q} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}} \right) \right) \delta q + \cancel{\frac{\partial L}{\partial \dot{q}} \delta q} \Big|_{t_1}^{t_2} \\
 &\Rightarrow \boxed{\frac{\partial L}{\partial q} - \frac{d}{dt} \frac{\partial L}{\partial \dot{q}} = 0} \quad \text{Euler-Lagrange (ELa) equation.}
 \end{aligned}$$

-) example: $L = \frac{1}{2} \dot{q}^2 - V(q) \Rightarrow \ddot{q} = -\frac{dV}{dq}$
 $V=0: \ddot{q}=0, q(t) = q_0 + vt$ (free particle moves on straight line)
 $\delta S \neq 0$. 

-) Similar in field theory, except that we replace $q(t)$ by a set of spacetime dependent fields $\phi^i(x^\mu)$.
-) Since the field now depends on x^μ and not just t , the Lagrangian L can be expressed as spatial integral over a Lagrangian density $\mathcal{L}$.

$$L = \int d^3x \mathcal{L}(\phi^i, \partial_\mu \phi^i).$$

-) The action is then again the time integral of L

$$S = \int dt L = \int d^4x \mathcal{L}(\phi^i, \partial_\mu \phi^i).$$

-) The ELG equations follow then again from demanding that small variations of the fields leave the action unchanged

$$\phi^i \rightarrow \phi^i + \delta\phi^i, \quad \partial_\mu \phi^i \rightarrow \partial_\mu \phi^i + \delta\partial_\mu \phi^i = \partial_\mu \phi^i + \partial_\mu \delta\phi^i,$$

while keeping its boundary value fixed: $\delta\phi^i|_{\text{bdry}} = 0$.

-) Since $\delta\phi^i$ is by definition small, we can Taylor expand the Lagrangian:

$$\mathcal{L}(\phi^i, \partial_\mu \phi) \rightarrow \mathcal{L}(\phi^i + \delta\phi^i, \partial_\mu \phi^i + \partial_\mu \delta\phi^i),$$

$$= \mathcal{L}(\phi^i, \partial_\mu \phi) + \frac{\partial \mathcal{L}}{\partial \phi^i} \delta\phi^i + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \partial_\mu (\delta\phi^i),$$

$$\Rightarrow S \rightarrow S + \delta S, \quad \delta S = \int d^4x \left(\frac{\partial \mathcal{L}}{\partial \phi^i} \delta\phi^i + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \partial_\mu (\delta\phi^i) \right).$$

-) Now we would like to factor out $\delta\phi^i$ like before

$$\int d^4x \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \partial_\mu \delta\phi^i = \int d^4x \left(\partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \delta\phi^i \right) - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \right) \delta\phi^i \right)$$

↓
total derivative

$$\delta\phi^i|_{\text{bdry}} \stackrel{!}{=} 0 \Rightarrow \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \delta\phi^i|_{\text{bdry}} = 0$$

$$\Rightarrow \delta S = \int d^4x \left(\frac{\partial \mathcal{L}}{\partial \phi^i} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \right) \right) \delta\phi^i.$$

$$\text{ELG: } \boxed{\frac{\partial \mathcal{L}}{\partial \phi^i} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi^i)} \right) = 0.}$$

•) Example 1: Real scalar field Lorentz scalar

$$\mathcal{L} = -\frac{1}{2}(\partial\phi)^2 - V(\phi) = -\frac{1}{2}\eta^{\mu\nu}\partial_\mu\phi\partial_\nu\phi - V(\phi)$$

$$= \underbrace{\frac{1}{2}\dot{\phi}^2}_{\text{kinetic energy}} - \underbrace{\frac{1}{2}(\vec{\nabla}\phi)^2}_{\text{gradient energy}} - \underbrace{V(\phi)}_{\text{potential energy}} \quad \text{"} \mathcal{L} = K - V \text{"}$$

$$\text{ELG: } 0 = -\frac{\partial\mathcal{L}}{\partial\phi} + \partial_\mu\frac{\partial\mathcal{L}}{\partial(\partial_\mu\phi)}$$

$$\text{first term is easy: } \frac{\partial\mathcal{L}}{\partial\phi} = -\frac{dV}{d\phi}$$

second term is a bit more tricky so we do it step-by-step:

$$\eta^{\mu\nu}\partial_\mu\phi\partial_\nu\phi = \eta^{\rho\sigma}\partial_\rho\phi\partial_\sigma\phi \quad | \text{relabel dummy indices}$$

$$\text{now we can use the general rule: } \frac{\partial V_a}{\partial V_b} = \delta_a^b$$

$$\frac{\partial}{\partial(\partial_\mu\phi)}(\eta^{\rho\sigma}\partial_\rho\phi\partial_\sigma\phi) = \eta^{\rho\sigma}(\delta_\rho^\mu(\partial_\sigma\phi) + (\partial_\rho\phi)\delta_\sigma^\mu)$$

$$= \eta^{\mu\sigma}\partial_\sigma\phi + \eta^{\rho\mu}\partial_\rho\phi = \eta^{\mu\nu}\partial_\nu\phi + \eta^{\nu\mu}\partial_\nu\phi = 2\eta^{\mu\nu}\partial_\nu\phi$$

putting the two terms together:

$$-\partial_\mu(-\eta^{\mu\nu}\partial_\nu\phi) - \frac{dV}{d\phi} = 0$$

$$\eta^{\mu\nu}\partial_\mu\partial_\nu\phi - \frac{dV}{d\phi} = 0$$

d'Alembertian $\square$

$$\square\phi - \frac{dV}{d\phi} = 0$$

-) So far we did not assume a specific form of the metric.
On flat space $\eta_{\mu\nu}$ is Minkowski metric the equation reads:

$$\ddot{\phi} - \nabla^2 \phi + \frac{dV}{d\phi} = 0.$$

-) Assuming the potential of a harmonic oscillator $V(\phi) = \frac{1}{2} m^2 \phi^2$, the equation of motion reduces to the Klein-Gordon equation:

$$\square \phi - m^2 \phi = 0.$$

-) Example 2: $U(1)$ gauge field A_μ (electromagnetism)

Dynamical field is 4-vector potential
└ electrostatic potential

$$A_\mu = (\phi, \vec{A})$$

└ vector potential

The action can be expressed in terms of the field strength tensor:

$$F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu.$$

From this we see that $F_{\mu\nu}$ is invariant under gauge transformation

$$A_\mu \rightarrow A_\mu + \partial_\mu \lambda(x),$$

$$F_{\mu\nu} \rightarrow F_{\mu\nu} + \partial_\mu \partial_\nu \lambda - \partial_\nu \partial_\mu \lambda = F_{\mu\nu}.$$

Gauge invariance is a fundamental property of electromagnetism, so the action must be gauge and Lorentz invariant:

$$\mathcal{L} = -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + A_\mu \overset{\sim \text{source term}}{J}^\mu.$$

•) The ELG equations take the form

$$\frac{\partial \mathcal{L}}{\partial A_0} - \partial_\mu \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu A_0)} \right) = 0 ,$$

where the first term is again easy: $\frac{\partial \mathcal{L}}{\partial A_0} = \delta^\mu_0 J_\mu = J^0$.

Second term is again more tricky. Since we vary w.r.t. $\partial_\mu A_0$, we first rewrite F^2 as follows:

$$F_{\mu\nu} F^{\mu\nu} = F_{\alpha\beta} F^{\alpha\beta} = \eta_{\alpha\beta} \eta^{\beta\sigma} F_{\alpha\beta} F_{\rho\sigma} ,$$

$$\frac{\partial F_{\alpha\beta} F^{\alpha\beta}}{\partial (\partial_\mu A_0)} = \eta^{\alpha\beta} \eta^{\beta\sigma} \left(\frac{\partial F_{\alpha\beta}}{\partial (\partial_\mu A_0)} F_{\rho\sigma} + F_{\alpha\beta} \frac{\partial F_{\rho\sigma}}{\partial (\partial_\mu A_0)} \right) ,$$

$$\text{since } F_{\alpha\beta} = \partial_\alpha A_\beta - \partial_\beta A_\alpha .$$

$$\frac{\partial F_{\alpha\beta}}{\partial (\partial_\mu A_0)} = \delta^\mu_\alpha \delta^\beta_0 - \delta^\mu_\beta \delta^\alpha_0 , \text{ similar for } \frac{\partial F_{\rho\sigma}}{\partial (\partial_\mu A_0)} .$$

$$\Rightarrow \frac{\partial F^2}{\partial (\partial_\mu A_0)} = \eta^{\alpha\beta} \eta^{\beta\sigma} \left((\delta^\mu_\alpha \delta^\beta_0 - \delta^\mu_\beta \delta^\alpha_0) F_{\rho\sigma} + (\delta^\mu_\rho \delta^\sigma_0 - \delta^\mu_\sigma \delta^\rho_0) F_{\alpha\beta} \right) ,$$

$$= (\eta^{\mu\rho} \eta^{\sigma 0} - \eta^{\sigma\rho} \eta^{\mu 0}) F_{\rho\sigma} + (\eta^{\alpha\mu} \eta^{\beta 0} - \eta^{\alpha 0} \eta^{\beta\mu}) F_{\alpha\beta} ,$$

$$= F^{\mu 0} - F^{0\mu} + F^{\mu 0} - F^{0\mu} = 4F^{\mu 0} ,$$

$$\Rightarrow \frac{\partial \mathcal{L}}{\partial (\partial_\mu A_0)} = -F^{\mu 0} \Rightarrow \partial_\mu F^{0\mu} = J^0 .$$